Particle Physics in the Early Universe (FS 17)

EXERCISE SERIES 1 Dr. Achilleas Lazopoulos

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Exercise 1

To Einstein's equations as quickly as possible (following A. Zee: Gravity in a Nutshell)

(i) General Relativity postulates a connection between matter (or energy) that fills space and the geometry of that space. The metric $g_{\mu\nu}$ is seen as a dynamical quantity that changes from point to point in space-time, according to the corresponding equations of motion. We want to start from a field theoretical approach here, according to which, the equations of motion can be derived from minimizing an action, i.e. the integral over space-time of some Lagrangian density, with respect to varying the dynamical field, i.e. the metric. Let's begin from the action of some field as seen in Quantum Field Theory

$$I = \int d^4x \mathcal{L}(\phi(x)) \tag{1}$$

However, in curved spacetime, one can choose any kind of coordinates to describe it. Hence the action should be invariant under a general coordinate transformation

$$x^{\mu} \to x^{\prime \mu} = S^{\mu}_{\nu} x^{\mu} \tag{2}$$

The Lagranigian density $\mathcal{L}(\phi(x))$ should be a scalar under general coordinate transformations. The same, then, should be true for the integral measure. Show that the measure d^4x is not invariant under the transformation in eq.2. Show, however, that the quantity

$$d^4x\sqrt{-g}$$
 , $g \equiv det(g_{\mu\nu})$ (3)

is invariant. Hence we should look for an action of the form

$$I = \int d^4x \sqrt{-g} \mathcal{L}(\phi(x)) \tag{4}$$

(ii) Postulate that the Lagrangian is a scalar function of the 'field', i.e. the metric, and it should contain a term with a two derivatives with respect to the metric, in the same way that the Lagrangian of a field has a 'kinetic' term with two derivatives. (iii) Gravity is related to curvature, and curvature expresses the failure of commutativity in differentiation. So let's start with a covariant vector B_{ρ} . Show that the derivative of this object, $\vartheta_{\mu}B_{\rho}$ does not transform as a second rank tensor, under eq. 2. Instead, show that we have to define a new derivative, the covariant derivative,

$$D_{\mu}B_{\rho} \equiv \vartheta_{\mu}B_{\rho} - \Gamma^{\sigma}_{\mu\rho}B_{\sigma} \tag{5}$$

and fix the transformation properties of the object $\Gamma^{\sigma}_{\mu\rho}$ such $D_{\mu}B_{\rho}$ transforms as a second rank covariant tensor:

$$\Gamma_{\mu\rho}^{\prime\sigma} = (S^{-1})^{\beta}_{\mu} (S^{-1})^{\gamma}_{\rho} S^{\sigma}_{\alpha} \Gamma^{\alpha}_{\beta\gamma} + S^{\sigma}_{\alpha} \frac{\vartheta^2 x^{\sigma}}{\vartheta x^{\prime\mu} \vartheta x^{\prime\rho}} \tag{6}$$

Note that the covariant derivative defined here is completely analogous to the covariant derivative defined for gauge fields.

(iv) The non-commutativity of derivatives is expressed by defining the Riemann tensor $R^{\sigma}_{\mu\nu\rho}$ from

$$[D_{\mu}, D_{\nu}] B_{\rho} \equiv -R^{\sigma}_{\rho\mu\nu} B_{\sigma} \tag{7}$$

Show that

$$R^{\sigma}_{\rho\mu\nu} = \left(\vartheta_{\mu}\Gamma^{\sigma}_{\nu\rho} + \Gamma^{\sigma}_{\mu\kappa}\Gamma^{\kappa}_{\nu\rho}\right) - (\mu\leftrightarrow\nu) \tag{8}$$

Taking into account that (we don't show this here)

$$\Gamma^{\mu}_{\nu\rho} = \frac{1}{2} g^{\mu\tau} (\vartheta_{\mu} g_{\nu\tau} + \vartheta_{\nu} g_{\rho\tau} - \vartheta_{\tau} g_{\nu\rho}) \tag{9}$$

note that $R^{\sigma}_{\rho\mu\nu}$ contains terms with two derivatives, so it's the object we are looking for.

(v) Show that $R^{\lambda}_{\rho\mu\nu} = -R^{\lambda}_{\rho\nu\mu}$, $R_{\tau\rho\mu\nu} = -R_{\rho\tau\nu\mu}$, $R_{\tau\rho\mu\nu} = -R_{\mu\nu\tau\rho}$. Hence show that out of the Riemann tensor you can only construct one rank-2 tensor, the Ricci tensor

$$R_{\mu\nu} \equiv g^{\tau\rho} R_{\tau\mu\rho\nu} \tag{10}$$

and one scalar

$$R \equiv g^{\mu\nu} R_{\mu\nu} \tag{11}$$

(vi) We are thus led to the action (called the Einstein-Hilbert action) for empty space-time

$$I_{EH} = \int d^4x \sqrt{-g} CR(x) \tag{12}$$

Vary the action with respect to the metric, $g_{\mu\nu}$ to get Einstein's equation in empty space:

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 0 \tag{13}$$

Note: when you vary the action and you require $\delta I = 0$, there are three terms inside the integral:

$$\delta I_{EH} = \int d^4x \delta(\sqrt{-g}) C g^{\mu\nu} R_{\mu\nu} + \int d^4x \sqrt{-g} C \delta(g^{\mu\nu}) R_{\mu\nu} + \int d^4x \sqrt{-g} C g^{\mu\nu} \delta(R_{\mu\nu})$$
(14)

The last term is in fact a total derivative, so it does not contribute to the action. If you want to derive this, look at section VI.5 of Zee's book.

Note: You will need the identity $\delta\sqrt{-g} = \sqrt{-g}\frac{1}{2}g^{\mu\nu}\delta g_{\mu\nu}$. If you want to prove this, you need to use $\log(\det M) = \operatorname{Tr}(\log M)$, for $g_{\mu\nu}$ in the place of the matrix M.